## **Some Physical Applications of Weak Turbulent Theory**

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#### Kolmogorov-Obukhov spectrum

If the fluid is stirred by external forces at scale of order L, the spectral peak of the excited turbulence is posed at  $k \sim k_0 \sim 1/L$ . But right behind the peak the spectrum has the universal form:

$$I_k = \langle |V_k|^2 \rangle \simeq \alpha P^{2/3} k^{-11/3}$$

Here P is the flux of energy to high wave numbers, defined by an intensity of stirring. This is a famous Kolmogorov-Obukhov spectrum. If  $P \gg k_0^4 \nu^3$ , the Kolmogorov-Obukhov spectrum is realized inside the "interval of universality"

$$k_0 < k < k_m, \qquad k_m \simeq \left(\frac{P}{\nu^3}\right)^{1/4}$$

Kraichnan [1967]: The inverse cascade takes place in 2D turbulence of ideal incompressible fluid.

$$K = \frac{1}{2} \int |curlV|^2 d\vec{r}, \qquad curlV = \frac{\partial V_x}{\partial V_y} - \frac{\partial V_y}{\partial V_x}$$

In the area of small scales  $k > k_0$  another Kolmogorov-type spectrum is established:

$$I_k \simeq \alpha_q \frac{Q^{2/3}}{k^5}$$

Zakharov [1966]:

Direct Cascade: 
$$I_k = \alpha_p \frac{P^{1/3}}{k^{7/2}}, \quad \alpha_p = 0,219$$
  
Inverse Cascade:  $I_k = \alpha_q \frac{Q^{1/3}}{k^{10/3}}, \quad \alpha_q = 0.227$ 

#### **Turbulence of capillary waves: direct cascade of energy**

Capillary waves on the surface of deep incompressible fluid are described by the following Hamiltonian:

$$H = H_0 + H_1$$
  
$$H_0 = \int \omega_k a_k^* a_k dk, \quad H_1 = \frac{1}{2} \int V_{kk_1k_2} (a_k^* a_{k_1} a_{k_2} + a_k a_{k_1}^* a_{k_2}^*) \delta_{k-k_1-k_2} dk_1 dk_2 dk_3$$

Here  $a_k$  is the complex normal amplitude of capillary ripplons,  $\omega_k = (\sigma k^3)^{1/2}$  is the dispersion law of ripplons,  $\sigma$  is coefficient of surface tension,  $V_{kk_1k_2} \simeq \sigma^{1/4}k^{9/4}$  is a coupling coefficient of three-ripplon interaction. It is important that  $V_{kk_1k_2}$  is a homogenous function:

$$V_{\epsilon k,\epsilon k_1,\epsilon k_2} = \epsilon^{g/4} V_{kk_1k_2}$$

The rippion's wave action  $n_k$  is introduced as follow:

$$\langle a_k a_{k'}^* \rangle = n_k \delta_{k-k'}$$

Apparently  $n_k \dot{k}^2 \gg h$ , thus the problem is pure classical.  $n_k$  obeys kinetic equation

$$\frac{\partial n_k}{\partial t} = st(n_k), \qquad st(n_k) = \int [R_{kk_1k_2} - R_{k_1kk_2} - Rk_2kk_1]dk_1dk_2$$

$$R_{kk_1k_2} = 4\pi |V_{kk_1k_2}|^2 \delta_{k-k_1-k_2} \delta_{\omega_k-\omega_{k_1}-\omega_{k_2}} [n_{k_1}n_{k_2} - n_k n_{k_1} - n_k n_{k_2}]$$

Thereafter we assume that spectrum is isotropic, n = n(|k|), and introduce the energy spectrum:

$$\epsilon(\omega) = \omega n(\omega), \qquad n(\omega) = 2\pi k \frac{dk}{d\omega} n(\omega)$$

The isotropic kinetic equation can be written as follow:

$$\frac{\partial \epsilon_{\omega}}{\partial t} + \frac{\partial P}{\partial \omega} = 0$$

Here  $P = -2\pi \int_0^\omega \omega k \frac{dk}{d\omega} st d\omega$  is a flux of energy to high frequency region.

The stationary kinetic equation st = 0 is equivalent to  $\partial P/\partial \omega = 0$ ,  $P = P_0$ . If  $P_0 = 0$ , this equation has unique thermodynamic Rayley-Jean's solution:

$$n(\omega) = \frac{T}{\omega}$$

In a general case, it has two-parametric solution:

$$n(\omega) = P^{1/2} \omega^{-17/6} f\left(\frac{T}{P^{1/2}} \omega^{11/6}\right)$$

We must put T = 0 (the temperature of turbulence is zero!). Then we get:

 $n(\omega) = \alpha P^{1/3} \omega^{17/6}, \quad \alpha = f(0)$  is Kolmogorov constant

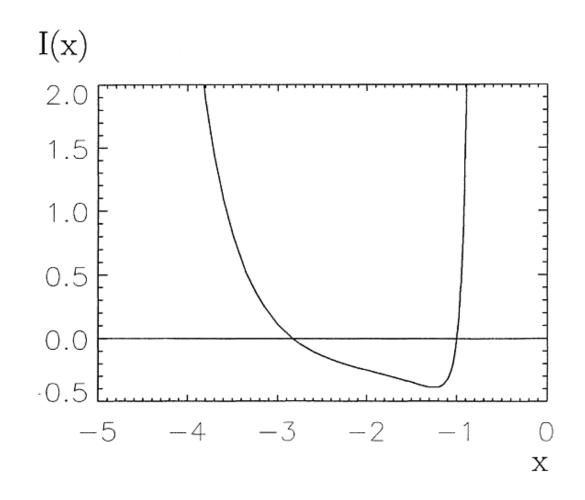
For the surface elevation spectrum  $I(\omega) = < |\omega|^2 >$  we get:

$$I(\omega) = \alpha \sigma^{2/3} \omega^{-19/6}$$

Existence of physically significant solution could be rigorously proven both analytically and numerically. For the numerical proof we can apply "frequency variables"  $\omega_i = \sqrt{\sigma k_i^3}$  and look for a solution in a powerlike form  $n(\omega) = A\omega^x$ . Then

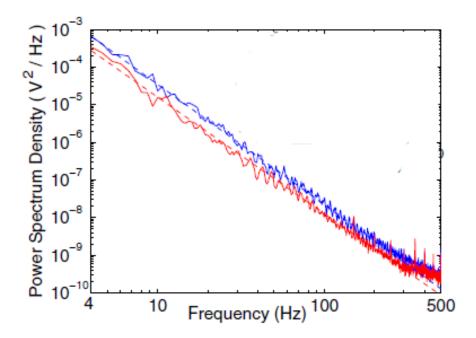
$$st = \frac{A^2 \omega^{-2x+7/6}}{36\pi \sigma^{5/3}} I(x)$$

The integrals in st are convergent if -5 < x < -5/6. Function I(x) is presented on next picture.



Collision integral I(x) as a function of solution index x: it is finite within the locality interval -5 < x < -5/6 and has two zeros at  $x_0 = -1$  (thermodynamical spectrum) and x = -17/6 (Kolmogorov solution).

KZ-spectrum for capillary waves was observed in many laboratory experiments. The most advanced one, performed in absence of gravity in free falling plane, showed that law  $I(\omega) \simeq \omega^{-19/6}$  was realized with a very good accuracy in the range of two decades.



Power spectrum density of surface wave height *in low gravity*. Lower curve: random forcing 0-6 Hz. Upper curve: sinusoidal forcing at 3 Hz. Dashed lines had slopes of -3.1 (lower) and -3.2 (upper).

### Inverse cascade in Bose-condensation

Weakly non-ideal Bose-gas (classical wave approximation, dimensionless variables):

 $i\Psi_t = \Delta \Psi - |\Psi|^2 \Psi$  - Nonlinear Schrodinger equation (NLSE)

Approximation of weak nonlinearity (in the absence of condensate):

 $<\Psi(k)\Psi^*(k')>=n_k\delta_{k-k'}$  – Particle distribution function for bose-particles

This function obeys kinetic equation:

$$\frac{\partial n}{\partial t} = st(n,n,n) =$$

The first task of weak turbulent theory - classification of solutions of equation:

$$st(n, n, n) = 0$$

We look only for isotropic solutions that depend on  $\omega = k^2$ . Moreover, we assume  $n(\omega) = A\omega^{-x}$ . After some transformations we get:

$$st = \pi^3 A^3 \omega^{-3x+2} F(x)$$

$$F(x) = \int_0^1 d\xi \int_1^\infty d\eta [1 + (\xi + \eta - 1)^x - \xi^x - \eta^x] [1 + (\xi + \eta - 1)^{3x - 7/2} - \xi^{3x - 7/2} - \eta^{3x - 7/2}]$$
  

$$F(x) \text{ has four solutions: } x = 0, x = 1, x = 7/6, x = 3/2. \text{ However, the integral converges only if}$$

Thus two solutions only make sense:

$$n_1(k) = \frac{T}{k^2}$$
 — This is thermodynamic Rayley-Jean spectrum

$$n_2(k) = \frac{\alpha Q^{1/3}}{k^{7/3}}$$
 — This is KZ spectrum for inverse cascade

Q is the flux of particles to small scales. Only these two spectra are "local"; two other are "nonlocal" and cannot be realized. Kinetic equation for particle distribution function has a family of self-similar solutions depending on arbitrary parameter  $\beta$ :

$$n \simeq \frac{1}{|t_0 - t|^{2\beta + 1/2}} F\left(\frac{k}{(t_0 - t)^{\beta}}\right)$$

In this case  $Q \simeq (t_0 - t)^{\beta - 3/2}$ . Compensation takes place if  $\beta > 0$ ,  $k_n \simeq (t_0 - t)^{\beta}$ . To define  $\beta$  we must know the rate of cooling of Bose-gas.

# Theory of wind-driven sea: coexistence of direct and inverse cascades

If  $\omega_k = \sqrt{gk}$  is dispersion relation of gravity waves on deep water, the lowest-order nonlinear process in the wind-driven sea is four-wave interaction governed by resonant condition:

$$\vec{k_1} + \vec{k_2} = \vec{k_3} + \vec{k_4}, \quad \omega_{k_1} + \omega_{k_2} = \omega_{k_3} + \omega_{k_4}$$

Gravity surface waves are described by Hamiltonian:

$$\int \omega_k b_k^* b_k dk + \frac{1}{4} \int T_{kk_1, k_2k_3} b_k^* b_{k_1}^* b_{k_2} b_{k_3} \delta_{k+k_1-k_2-k_3} dk dk_1 dk_2 dk_3$$

 $b_k$  is "renormalized" complex normal amplitude of gravitational ripplons.  $T_{kk_1k_2k_3}$ , the coupling coefficient of four-wave interaction, is a complicated homogenous function of third order:

$$T_{\epsilon k,\epsilon k_1,\epsilon k_2,\epsilon k_3} = \epsilon^3 T_{kk_1k_2k_3}$$

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We introduce the wave action spectrum of ripplons by standard way:

$$\langle b_k b_{k'}^* \rangle = n_k \delta_{k-k'}$$

A slightly inhomogeneous wind-driven sea  $n_k$  obeys the Hasselmann kinetic equation:

$$\begin{aligned} \frac{\partial n_k}{\partial t} &= S_{nl} + S_{in} + S_{diss} \\ S_{nl} &= F_k - n_k \Gamma_k \\ F_k &= \pi \int |T_{kk_1k_2k_3}|^2 n_{k_1} n_{k_2} n_{k_3} \delta_{k+k_1-k_2-k_3} b_{\omega_k+\omega_{k_1}-\omega_{k_2}-\omega_{k_3}} dk_1 dk_2 dk_3 \\ \Gamma_k &= \pi \int |T_{kk_1k_2k_3}|^2 (n) k_1 n_{k_2} + n_{k_1} n_{k_3} - n_{k_2} n_{k_3}) \delta_{k+k_1-k_2-k_3} \delta_{\omega_k+\omega_{k_1}-\omega_{k_2}-\omega_{k_3}} dk_1 dk_2 dk_3 \end{aligned}$$

 $\Gamma_k$  can be treated as an "effective dissipation" due to nonlinear four-wave interaction.

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The term  $S_{in}$  appears due to interaction with wind.

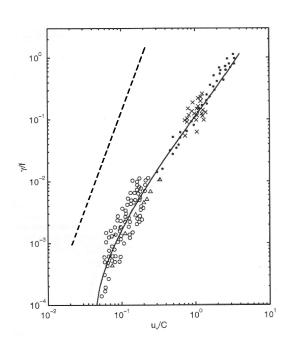
The term  $S_{diss} < 0$  is the rate of dissipation.

$$S_{in} + S_{diss} > 0$$

 $S_{in} = \gamma_k n_k, \quad \gamma_k$  is the growth rate of instability

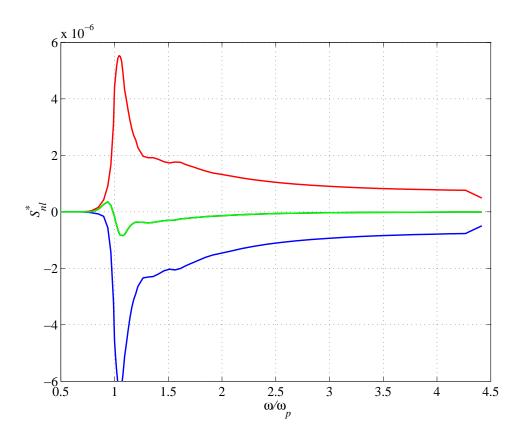
 $\Gamma_k$  surpasses  $\gamma_k$  at least in order of magnitude! The "conservative" equation:

$$\frac{\partial n_k}{\partial t} = S_{nl} = F_k - n_k \Gamma_k$$



Dots and solid line: experimental data for the wind-induced growth rate  $2\pi \gamma_{in}(\omega)/\omega$ . Dashed line: the damping due to four-wave interactions  $2\pi \Gamma(\omega)/\omega$ , calculated for narrow in angle spectrum at  $\mu \simeq 0.05$  using equation:

$$\Gamma_{\omega} = 36 \,\pi \omega \left(\frac{\omega}{\omega_p}\right)^3 \,\mu_p^4 \,\cos^2 \theta$$



Split of nonlinear interaction term  $S_{nl}$  (central curve) into  $F_k$  (upper curve) and  $\Gamma_k N_k$  (lower curve). For nonlinear interaction term  $S_{nl} = F_k - \Gamma_k N_k$  the magnitudes of constituents  $F_k$  and  $\Gamma_k N_k$  essentially exceed their difference. They are one order higher than the magnitude of  $S_{nl}$ !

The "conservative" equation

$$\frac{\partial n_k}{\partial t} = S_{nl} = F_k - n_k \Gamma_k$$

conserves two motion constants, energy and momentum:

$$E = \int \omega_k n_k dk \qquad \vec{P} = \int \vec{k} n_k dk$$

and the total wave action (or number of particles):

$$N = \int n_k d\vec{k}$$

In the isotropic situation  $\vec{P} = 0$ .

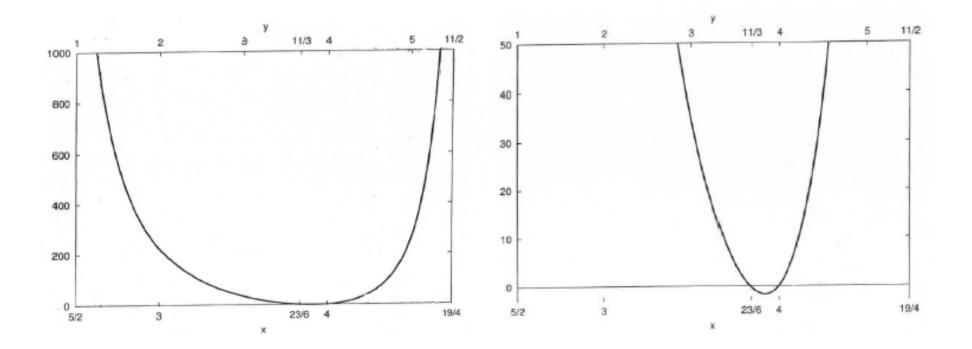
The central question of theory is classification of solutions of this equation:

$$S_{nl} = 0$$

The isotropic powerlike solutions  $N_k = k^{-x}$ .

$$S_{nl} = g^{3/2} k^{-3x+19/2} F(x)$$

The integral in  $S_{nl}$  converges if 3/2 < x < 19/4. F(x) = 0 if F(x) has only two zeros: x = 4 and x = 23/3. The numerical experiment confirms this fact with a great accuracy.



(a): Plot of function F(x).

(b): Plot of function F(x). Zoom in the vertical direction.

$$N_k^{(1)} = c_p \left(\frac{P_0}{g}\right)^{1/3} \frac{1}{k^4}$$
$$N_k^{(2)} = c_q \left(\frac{Q_0}{q^{3/2}}\right)^{1/3} \frac{1}{k^{23/6}}$$

In a typical situation the spectrum has a peak at  $k \sim k_0$  and the "rear face" at  $k > k_0$ . This is a field where inverse and direct cascades coexist and compete.

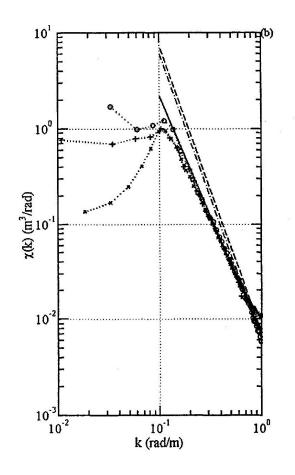
The angle energy spectrum:

$$\epsilon_k \simeq 2\pi k\omega_k N$$

For direct cascade:  $\epsilon_k \simeq k^{-5/2} \simeq k^{-2.5}$ 

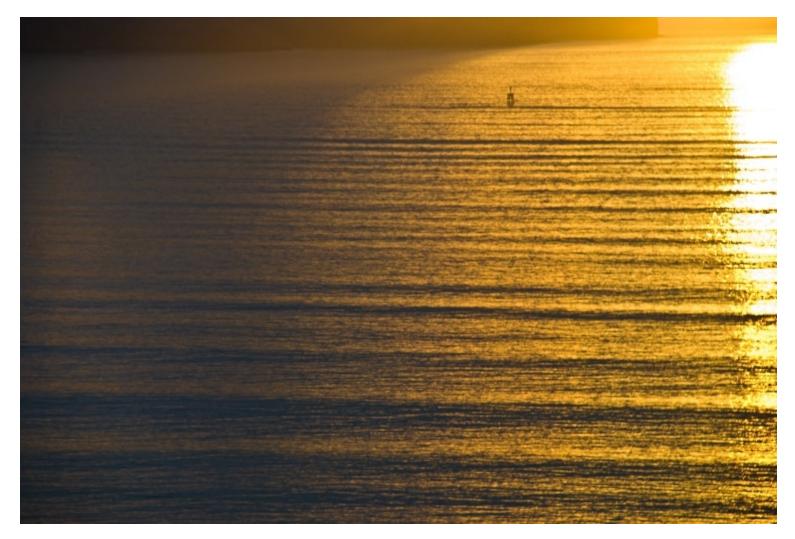
For inverse cascade:  $\epsilon_k \simeq k^{-7/3} \simeq k^{-2.33}$ 

In reality:  $\epsilon_k \simeq k^{-x}$ , where 2.33 < x < 2.5



Surface elevation spectra proportional to E(k)/g with comparison slopes. Solid line:  $k^{-5/2}$ ; dashed and dashed dotted curves:  $k^{-3}$  (Phillips's spectrum) with different normalizations.

## Waves forecasting.



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